

Sensitivity of the atmospheric response to the specification of the UV solar cycle variation in WACCM-3.5 simulations

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Observational and modeling evidence of quasi-decadal oscillations in the stratosphere and troposphere

- Solar cycle

Observational and modeling evidence of quasi-decadal oscillations in the stratosphere and troposphere

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- ▶ Mechanisms: top DOWN, bottom UP?

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Observational and modeling evidence of quasi-decadal oscillations in the stratosphere and troposphere

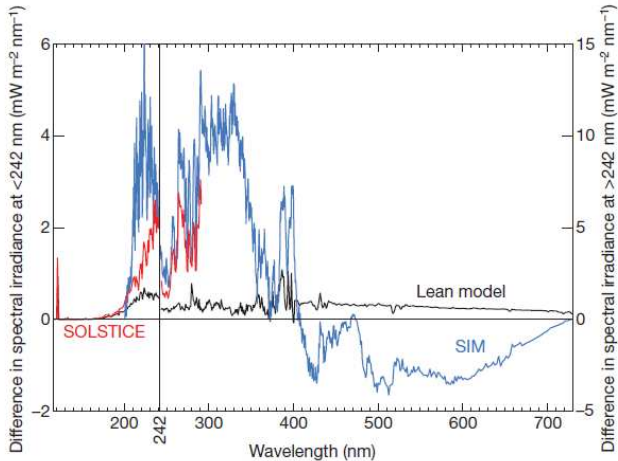
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The magnitude of 11-yr variations in SSI in SORCE-SIM retrievals is 5-7 times greater than NRL in the UV range, and out-of-phase in the VIS range (?)

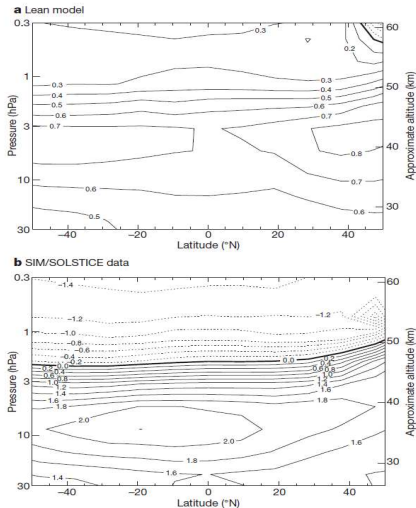
Some authors have investigated the response of climate models to solar-cycle variations in SSI derived from SORCE-SIM data

- Haigh et al., NATURE [2010]



Some authors have investigated the response of climate models to solar-cycle variations in SSI derived from SORCE-SIM data

- ▶ Haigh et al., NATURE [2010]
- ▶ Ozone response from Lean model and SORCE-SOLSTICE data



**Haigh et al. (2010) forced 2-D model with
SOLSTICE-SIM (2004-2007) spectra, including
out-of-phase VIS irradiance changes. Focus on O3
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Merkel et al. (2011) forced WACCM model with SOLSTICE-SIM (2004-2007) spectra. Focus on O3 response in the stratosphere and mesosphere

Ineson and Scaife (2011) used SIM-based (200-320nm) spectrum changes (2004-2007) in MetOffice model

Problem

SORCE-SIM data only represent changes over part of a solar cycle. The semi-empirical models (e.g. Lean 2005, Krivova 2006) reconstruct variations over multiple cycles.

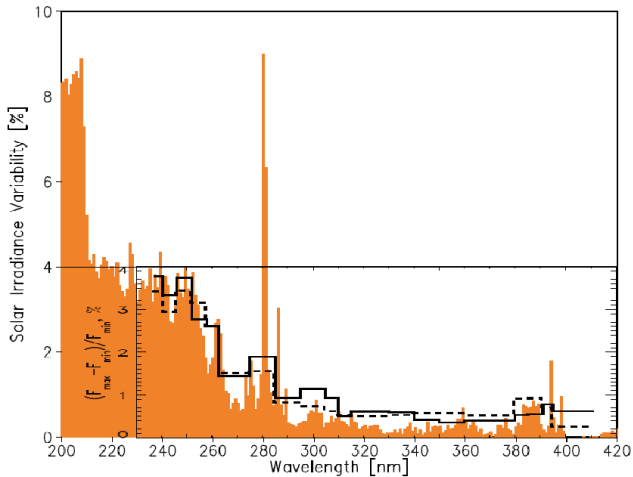


Figure: Relative 11-yr solar cycle variation in spectral solar irradiance modeled by Lean (2005) (orange bars) and Krivova (2006) (black line)

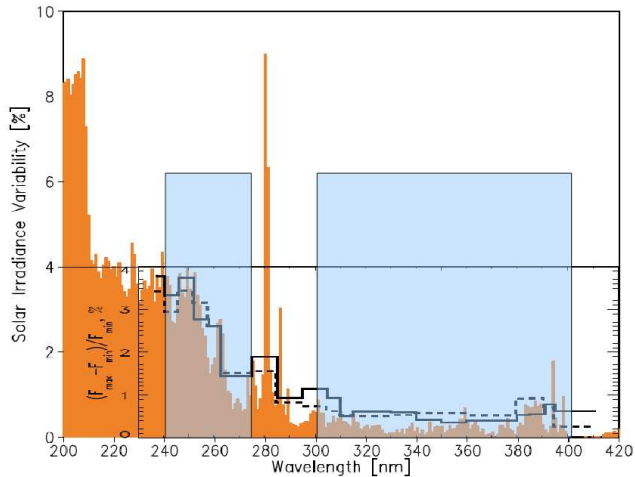
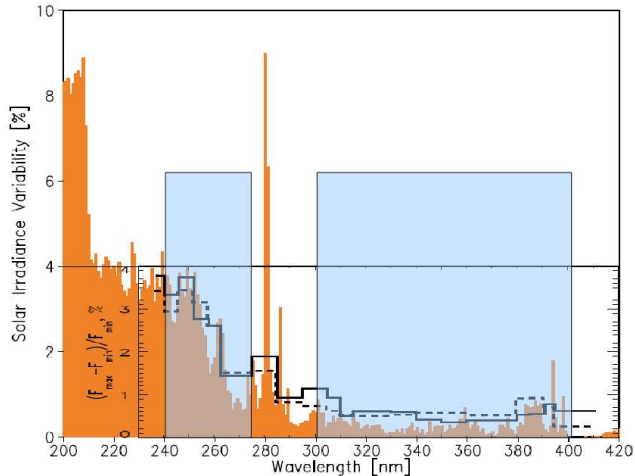


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TASK: Study the sensitivity of a GCM response to uncertainty in the SSI forcing.

Model and experimental set-up

Whole Atmosphere Community Climate Model - v3.5

Model features:

- ▶ Finite-Volume dynamical core (Lin Rood 2004)
- ▶ $1.9 \times 2.5 \times 66$ L (model top at 140 km)
- ▶ interactive chemistry (Kinnison et al., 2007)

Set-up:

- ▶ fixed solar spectral irradiance (MAX vs MIN spectra)
- ▶ climatological SSTs, no QBO, 1995-composition

Model and experimental set-up

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WACCM-3.5 reproduces most of the features of the
observed signal from the 11-year solar cycle

Model and experimental set-up

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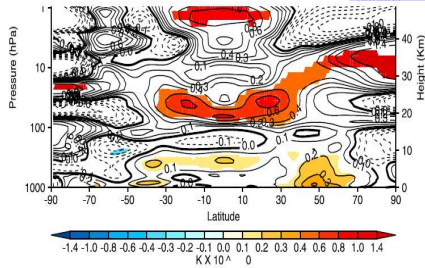
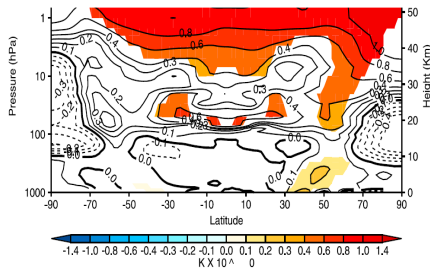


Figure: Annual mean solar (F10.7) regression coefficient in zonal mean temperature, from WACCM-3.5 (left) and ERA-40 data (right)

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Time-slice experiments

Experiment Name	solar forcing	length
NRL-sMIN	solar MINIMUM from Lean (2005), SC 21-24 ave	20 yy (1 spinup)
NRL-sMAX	solar MAXIMUM from Lean (2005), SC 21-24 ave	20 yy (1 spinup)
sMAX*[300-400]	solar MAXIMUM from Lean (2005) + 1%[300-400]	20 yy (1 spinup)
sMAX*[240-270]	solar MAXIMUM from Lean (2005) + 1%[240-270]	20 yy (1 spinup)

- ▶ "Solar cycle Lean": NRL-sMAX - NRL-sMIN
- ▶ "Solar cycle Lean + UV [300-400]": sMAX*[240-270] - NRL-sMIN

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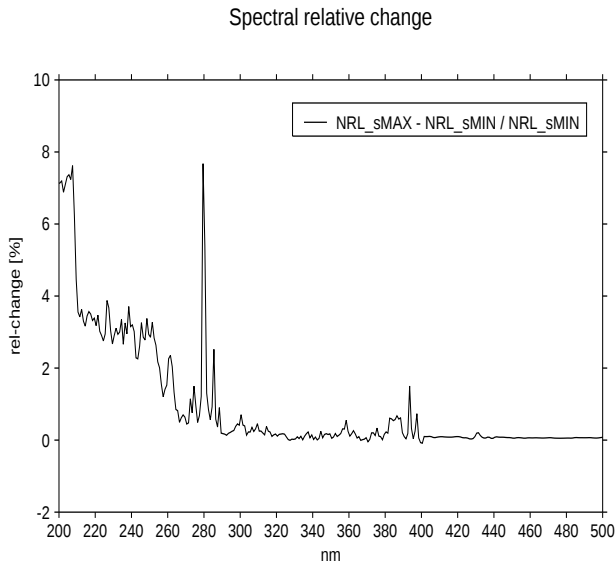
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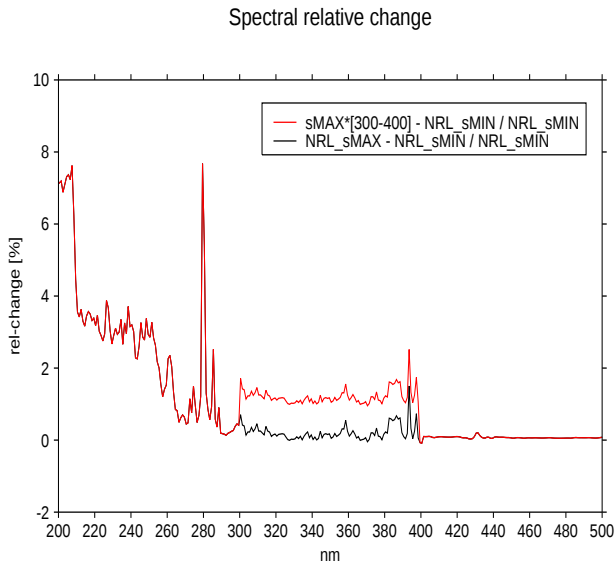
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Solar forcing input

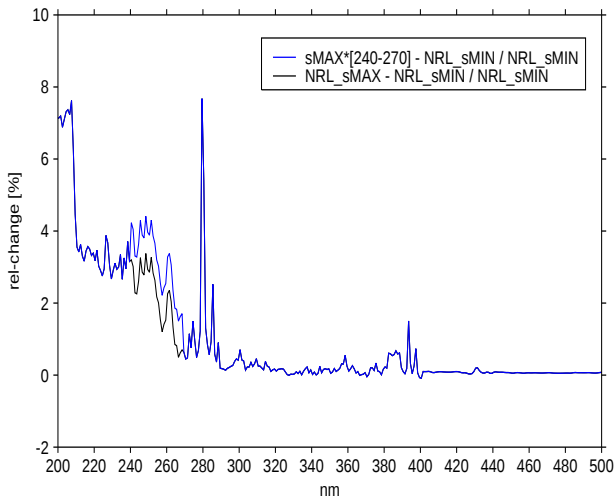


Solar forcing input



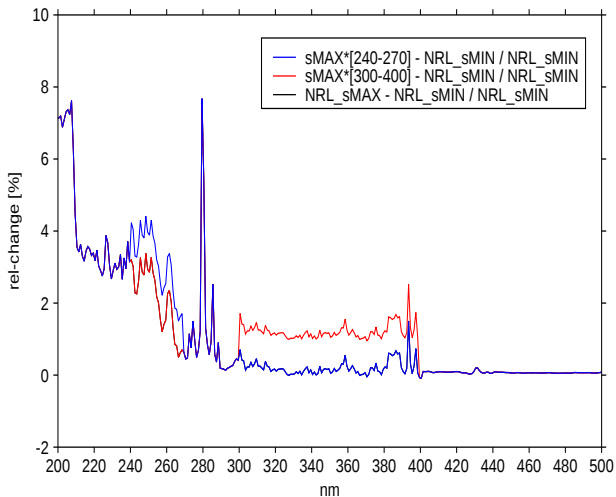
Solar forcing input

Spectral relative change



Solar forcing input

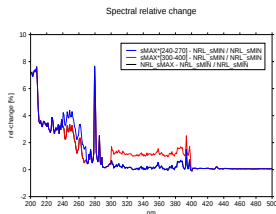
Spectral relative change



Solar forcing input

TABLE 1b. Summary of Spectral Data

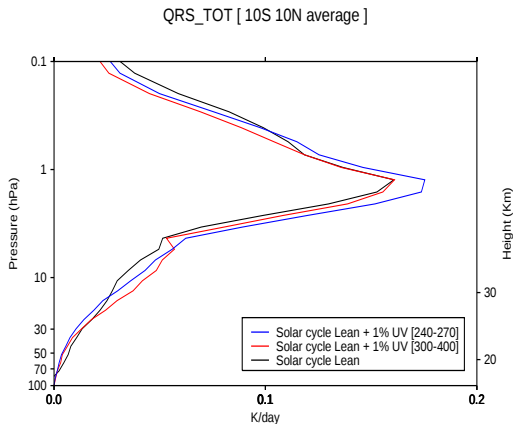
Interval	min ^a	max ^b	$p(\text{H}_2\text{O})^c$	$k(\text{H}_2\text{O})^d$	$k(\text{droplets})^e$	$p(\text{CO}_2)^f$
1	0.200	0.245	0	0	1	0
2	0.245	0.265	0	0	1	0
3	0.265	0.275	0	0	1	0
4	0.275	0.285	0	0	1	0
5	0.285	0.295	0	0	1	0
6	0.295	0.305	0	0	1	0
7	0.305	0.350	0	0	1	0
8	0.350	0.700	0	0	1	0
9	0.700	5.000	0.505	0.002	2	0
10	0.700	5.000	0.210	0.035	3	0
11	0.700	5.000	0.120	0.377	3	0
12	0.700	5.000	0.070	1.950	3	0
13	0.700	5.000	0.048	9.400	3	0
14	0.700	5.000	0.029	44.600	4	0
15	0.700	5.000	0.018	190.000	4	0
16	2.630	2.860	0	0	4	1.00
17	4.160	4.550	0	0	4	0.64
18	4.160	4.550	0	0	4	0.36



300-400 nm : 1 full band in radiation code (2 partial)

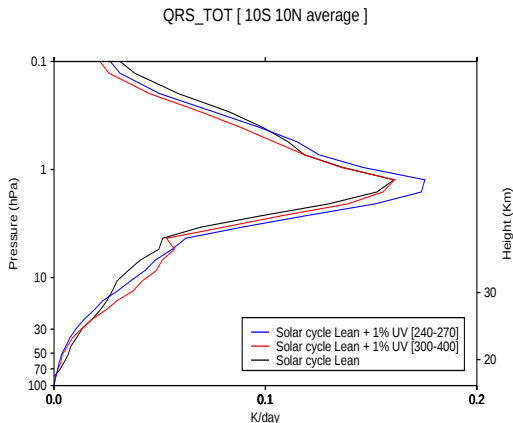
240-270 nm : 1 full band in radiation code (2 partial)

Radiative response



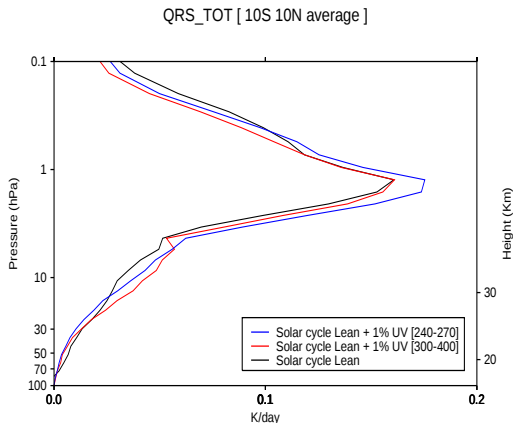
- Peak in SW-heating occurs at 45 km

Radiative response



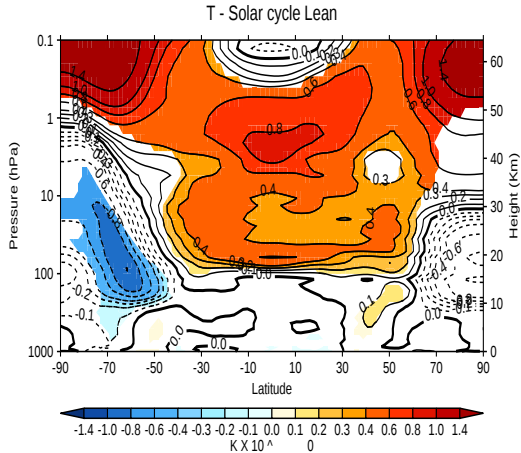
- ▶ Peak in SW-heating occurs at 45 km
- ▶ Slightly more SW-heating in the middle stratosphere if [300-400] radiation is enhanced

Radiative response



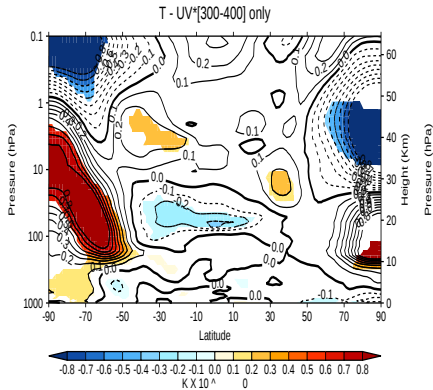
- ▶ Peak in SW-heating occurs at 45 km
- ▶ Slightly more SW-heating in the middle stratosphere if [300-400] radiation is enhanced
- ▶ More SW-heating in the middle and upper stratosphere if [240-270] radiation is enhanced

Solar cycle response in zonal mean temperature

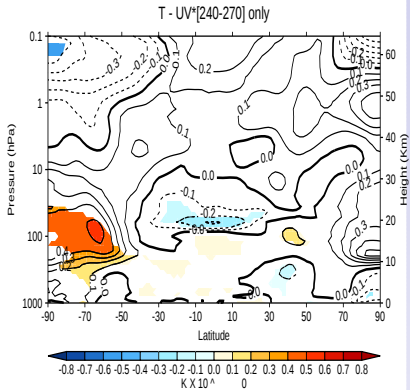


Impact of 1% change of solar radiation on temperature

SC Lean+UV[300-400]



SC Lean+UV[240-270]



Solar cycle response in zonal mean ozone

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sensitivity to UV
forcing

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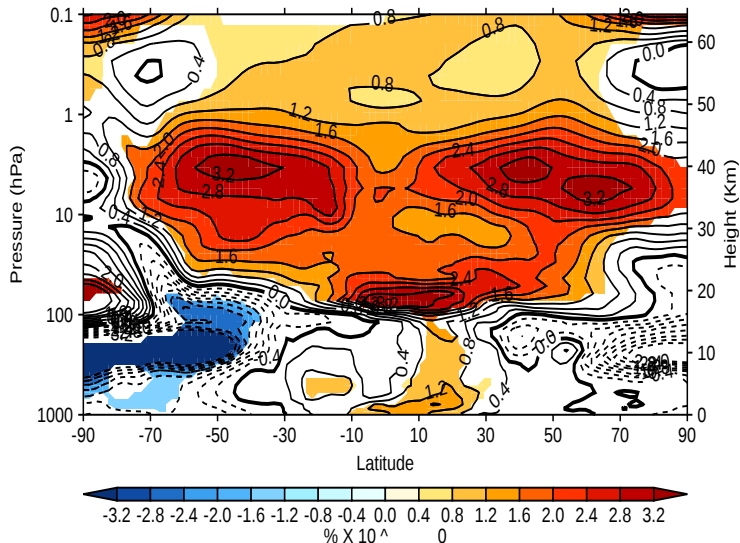
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O3 - Solar cycle Lean



Solar cycle response in zonal mean ozone

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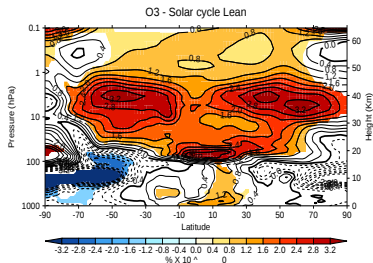
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Impact of 1% change of solar radiation on ozone

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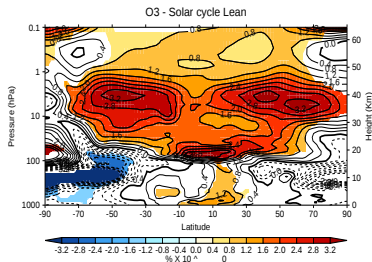
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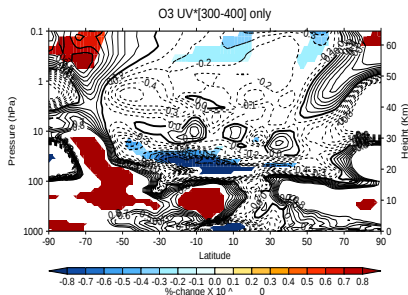
Results

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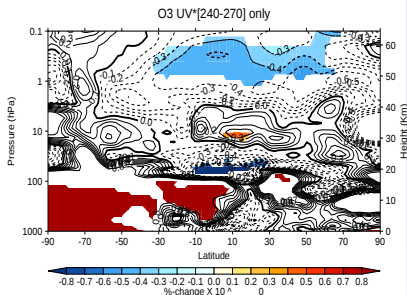
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UV + [300-400]



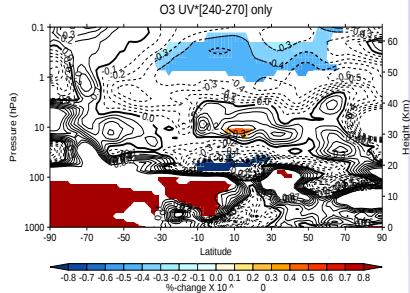
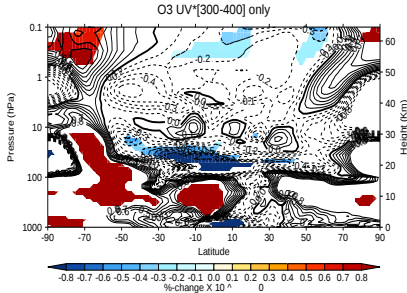
UV + [240-270]



Solar cycle response in zonal mean ozone

UV + [300-400]

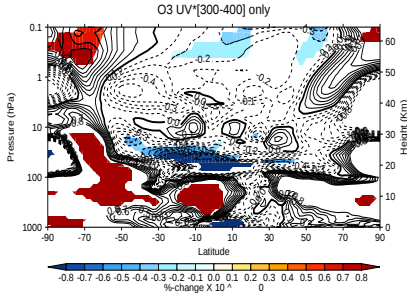
UV + [240-270]



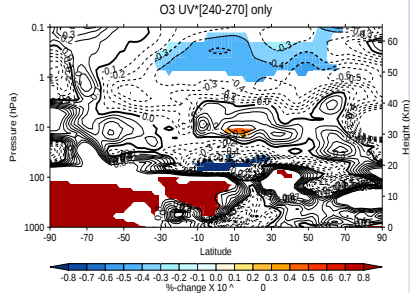
- Lower-tropical stratosphere (20 km): Less ozone if UV+ (240-270) and UV+ (300-400)

Solar cycle response in zonal mean ozone

UV + [300-400]



UV + [240-270]

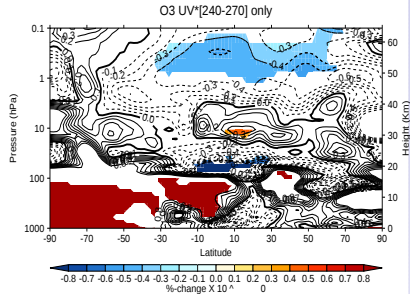
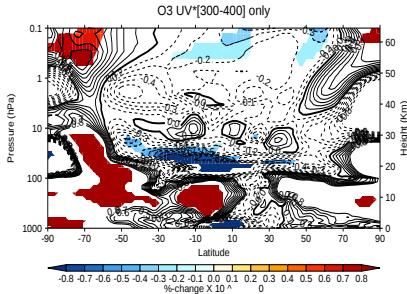


- ▶ Lower-tropical stratosphere (20 km): Less ozone if UV+ (240-270) and UV+ (300-400)
- ▶ Middle stratosphere (30 km): More ozone if UV+ (240-270)

Solar cycle response in zonal mean ozone

UV + [300-400]

UV + [240-270]



- ▶ Lower-tropical stratosphere (20 km): Less ozone if UV+ (240-270) and UV+ (300-400)
- ▶ Middle stratosphere (30 km): More ozone if UV+ (240-270)
- ▶ Upper stratosphere and lower mesosphere (40-50 km) : Less ozone if UV+ (240-270) and UV+ (300-400)
Less O3 possibly compensates more SW-absorption = less warming?

Seasonal solar cycle response in temperature

Model
sensitivity to UV
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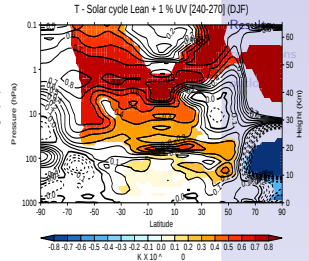
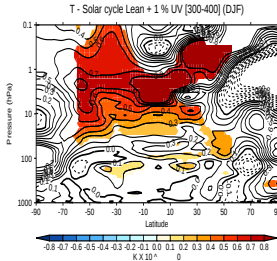
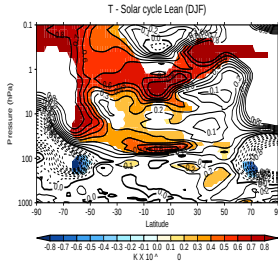
Boreal winter: DJF

SC

Lean

SC Lean + UV [300-400]

SC Lean + UV [240-270]



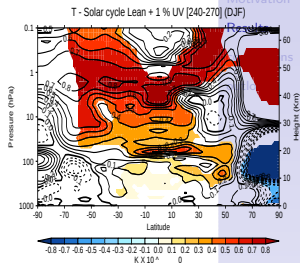
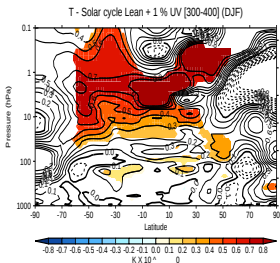
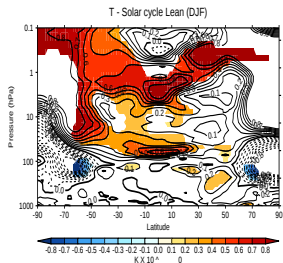
- Strong warming at 20 km over the equator is mostly produced during boreal winter in NRL-sMAX, along with NH polar cooling.

SC

Lean

SC Lean + UV [300-400]

SC Lean + UV [240-270]



- ▶ Strong warming at 20 km over the equator is mostly produced during boreal winter in NRL-sMAX, along with NH polar cooling.
- ▶ A similar (though stronger) pattern is also simulated if UV+ [240-270].

SC

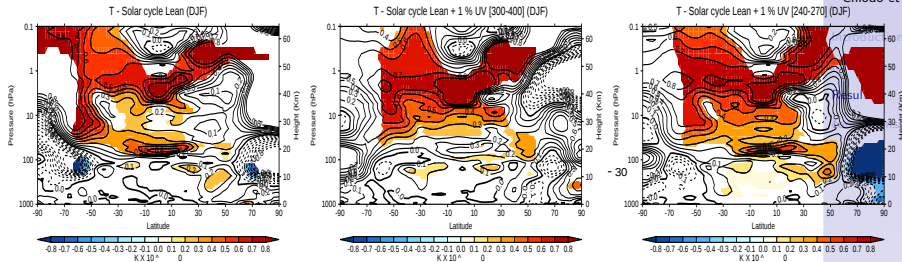
Lean

SC Lean + UV [300-400]

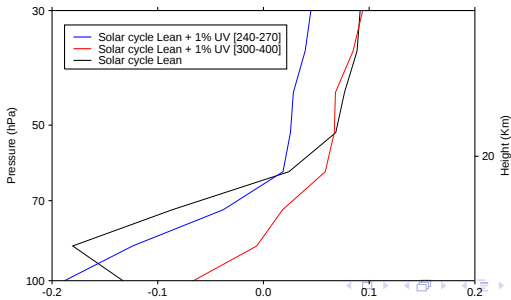
SC Lean + UV [240-270]

Model
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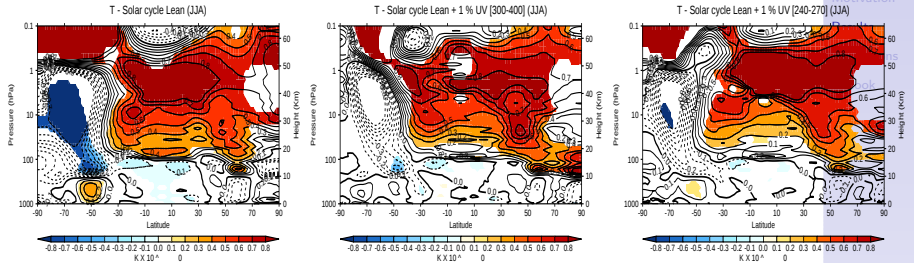
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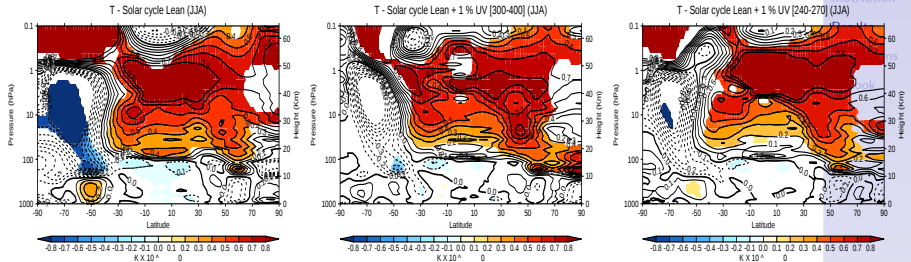
WSTAR [10S 10N] DJF



Austral winter: JJA



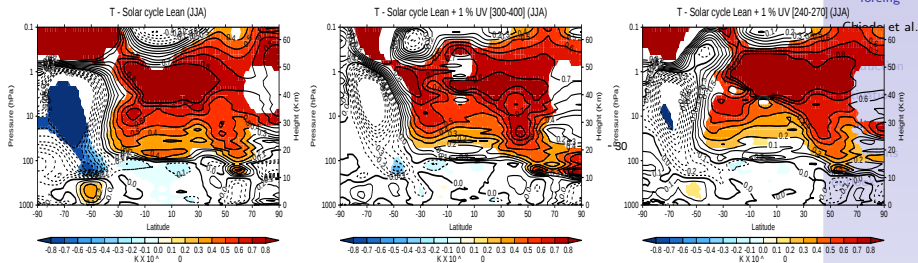
Austral winter: JJA



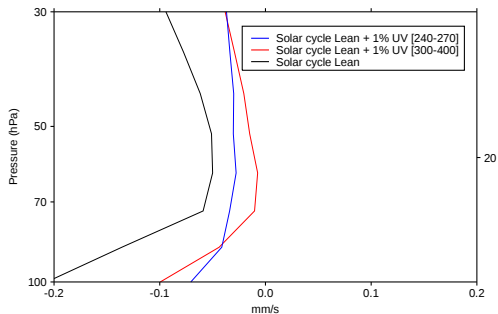
- A cooling at high and warming at low latitudes is simulated in the SH in NRL-sMAX. No such pattern is simulated if UV + [240-270] and UV + [300-400]

Austral winter: JJA

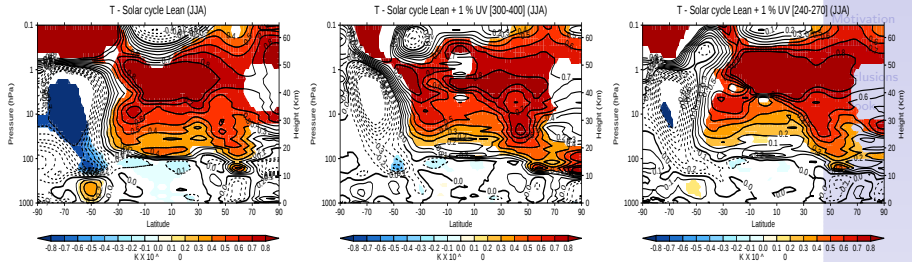
Model
sensitivity to UV
forcing



WSTAR [10S 10N] JJA



Austral winter: JJA



- Relative equatorial downwelling is simulated below 70 hPa in NRL-sMAX, and in UV + [240-270] compared to NRL-sMIN. Weakening of equatorial upward branch of B-Dobson circulation ?

Solar cycle response in zonal mean zonal wind in the NH

Model sensitivity to UV forcing

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U - Solar cycle Lean

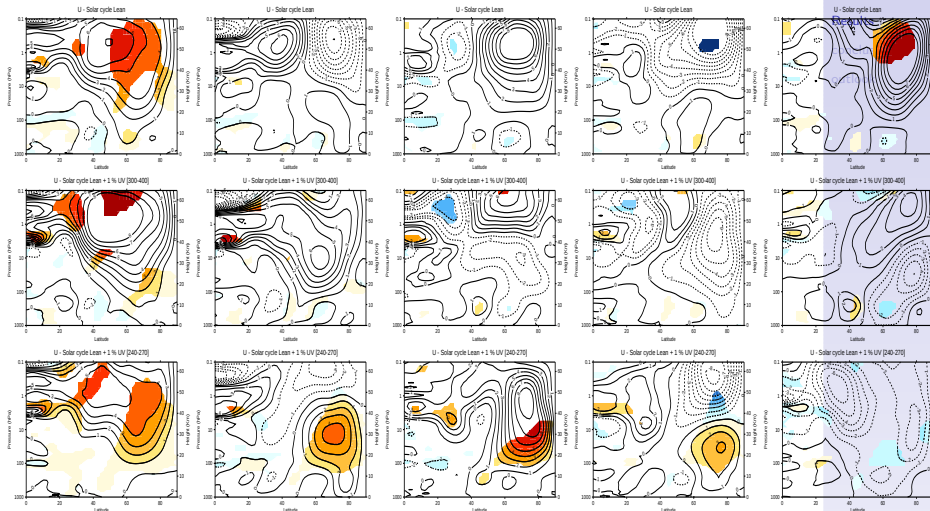
November

December

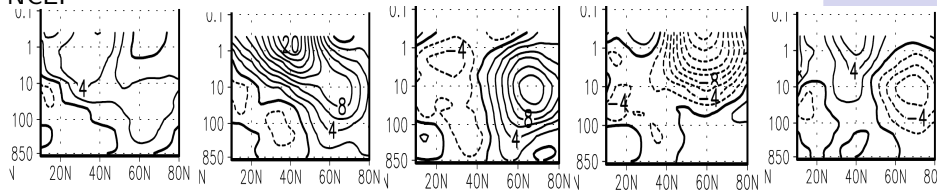
January

February

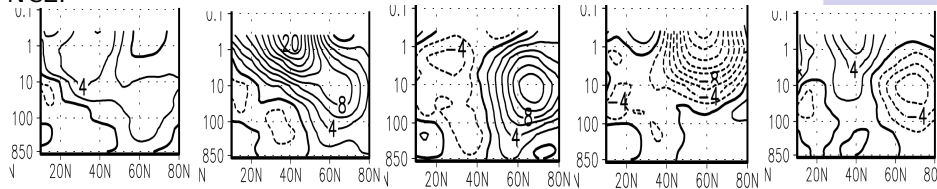
March



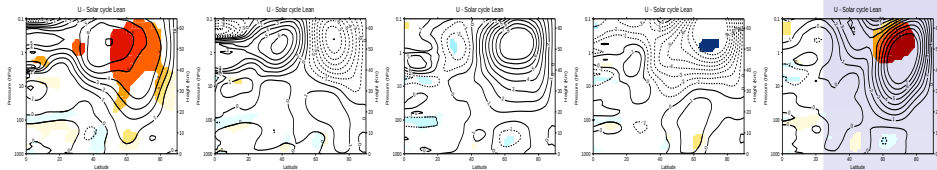
NCEP



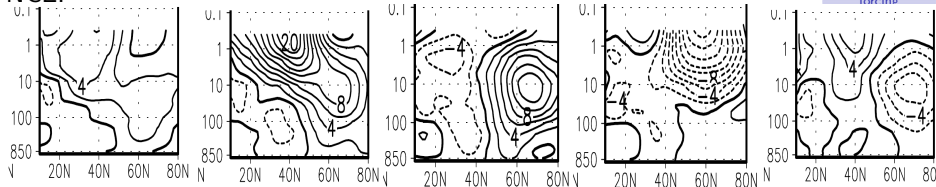
NCEP



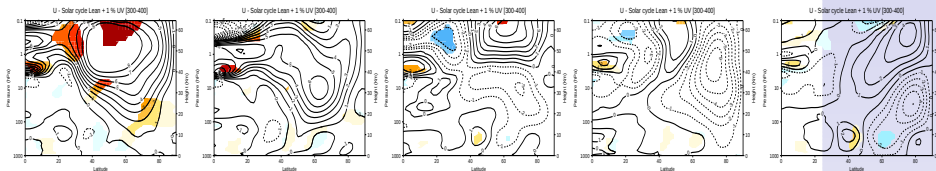
Solar cycle Lean



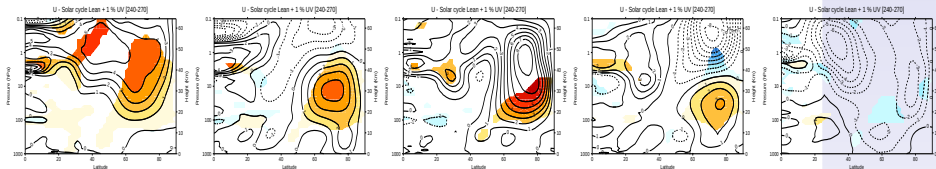
NCEP



Solar cycle Lean + UV [300-400]



Solar cycle Lean + UV [240-270]



Discussion

- ▶ The solar cycle derived from the Lean dataset (NRL-sMAX-NRL-sMIN) leads to stronger vortex in early-winter, but no clear response in the following months

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- ▶ The increase in the [300-400] radiation leads to stronger change in the Polar vortex during early winter. A switch in sign is found in January (i.e. Weaker vortex). More realistic late winter response

- ▶ The solar cycle derived from the Lean dataset (NRL-sMAX-NRL-sMIN) leads to stronger vortex in early-winter, but no clear response in the following months
- ▶ The increase in the [300-400] radiation leads to stronger change in the Polar vortex during early winter. A switch in sign is found in January (i.e. Weaker vortex). More realistic late winter response
- ▶ A stronger polar-vortex is simulated throughout the NH winter if the [240-270] radiation is enhanced. A switch in sign is found in March (i.e. Weaker vortex). More realistic early to mid winter response

Solar cycle response SW-heating rates from Lean (NRL-sMAX - NRL-sMIN) in November

Model
sensitivity to UV
forcing

Chiodo et al.

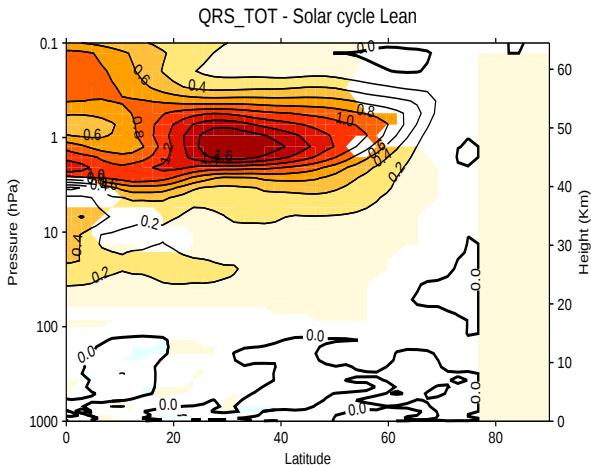
Introduction

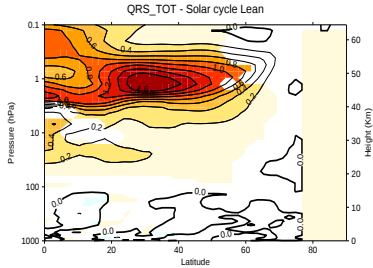
Motivation

Results

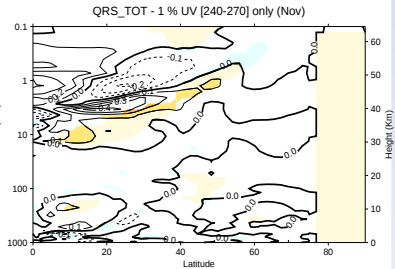
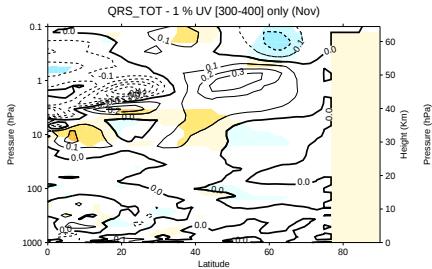
conclusions

outlook

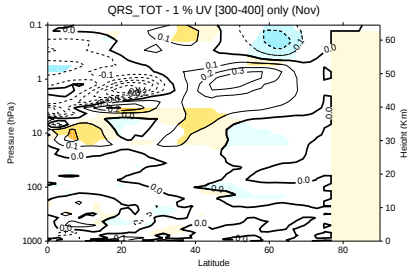




Impact of 1% radiation change in [300-400] and [240-270]



SC Lean+UV[300-400]



SC Lean+UV[240-270]

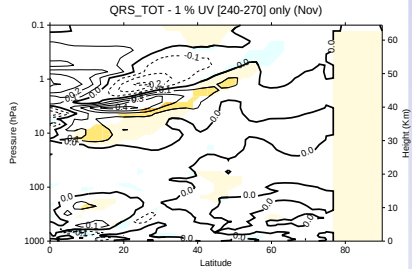
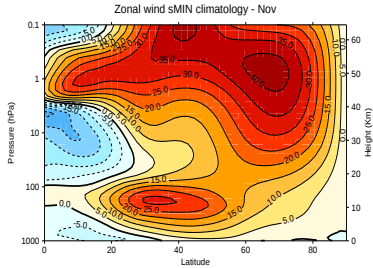
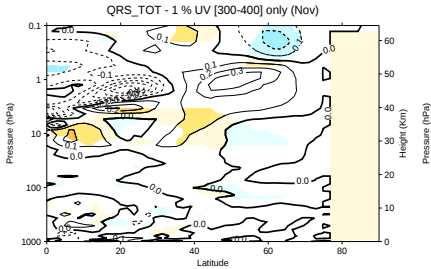


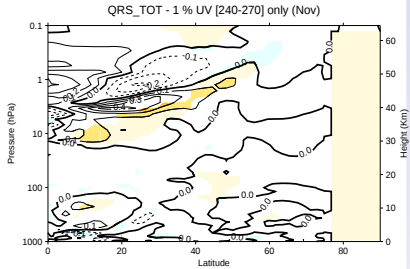
Figure: More SW-heating at 50-60N (left), and 30N (right)



SC Lean+UV[300-400]



SC Lean+UV[240-270]



Solar cycle response in Planetary-Wave divergence (EPF-div) from Lean (NRL-sMAX - NRL-sMIN) in November [m/s/day]

Model
sensitivity to UV
forcing

Chiodo et al.

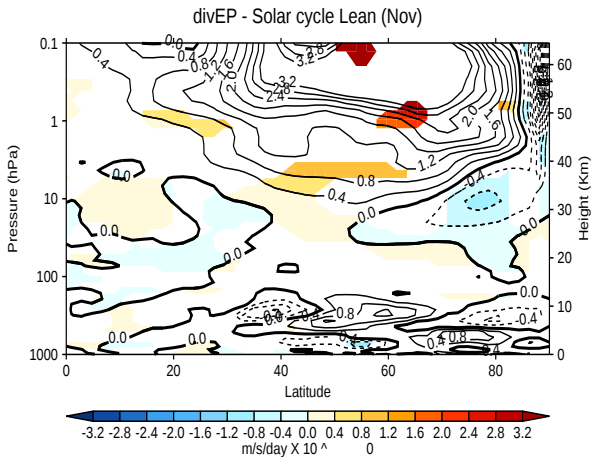
Introduction

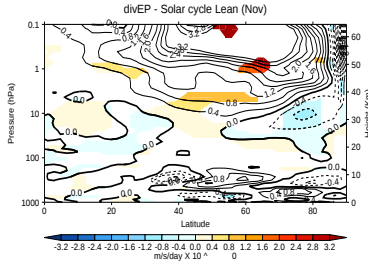
Motivation

Results

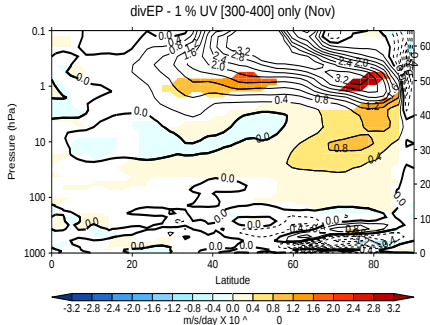
conclusions

outlook

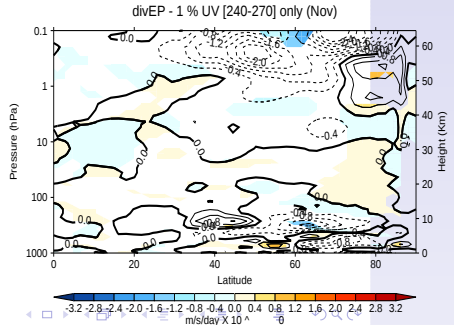




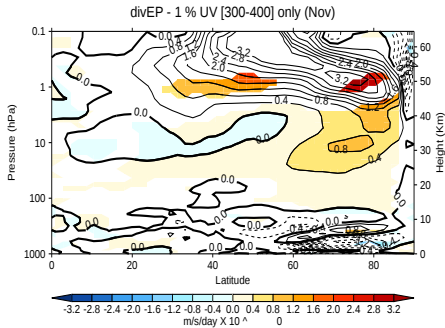
Impact of 1% [300-400]



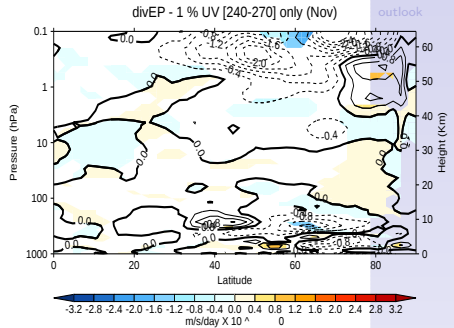
[240-270]



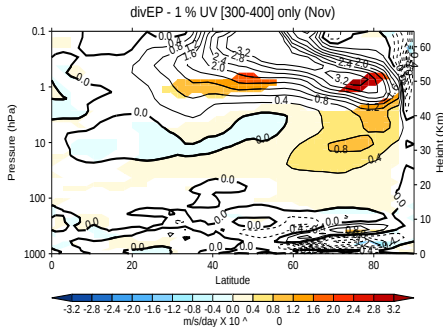
UV + [300-400]



UV + [240-270]



UV + [300-400]



UV + [240-270]

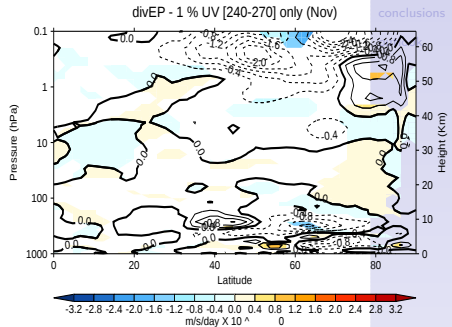


Figure: More divEP at stratopause (left), and less (right).

Planetary-Wave divergence changes in Dec and Jan

Model
sensitivity to UV
forcing

Chiodo et al.

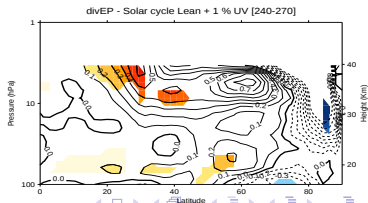
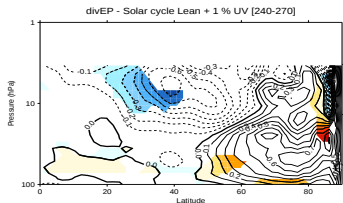
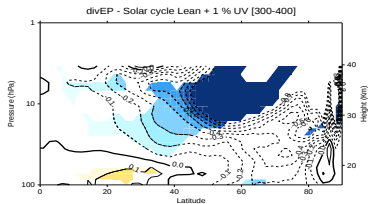
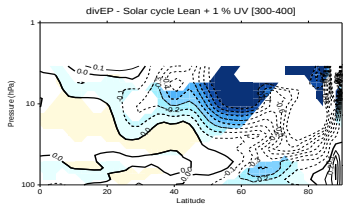
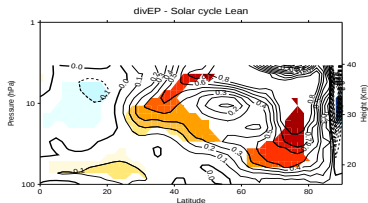
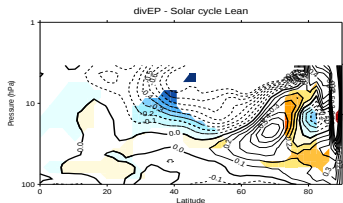
Introduction

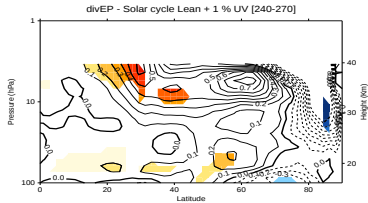
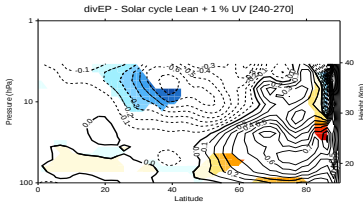
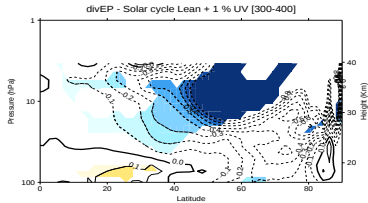
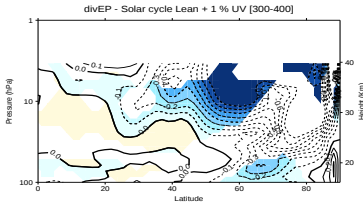
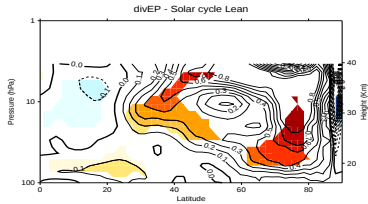
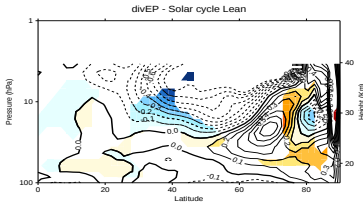
Motivation

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outlook





Residual circulation (v^* , w^*) changes in Dec and Jan

Model
sensitivity to UV
forcing

Chiodo et al.

Introduction

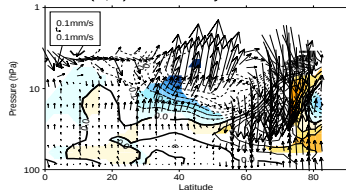
Motivation

Results

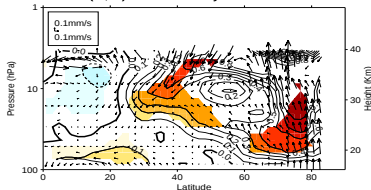
conclusions

outlook

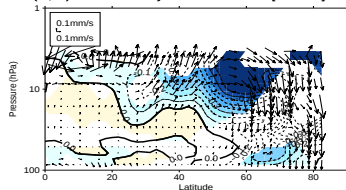
(v^* , w^*) divEP - Solar cycle Lean - Dec



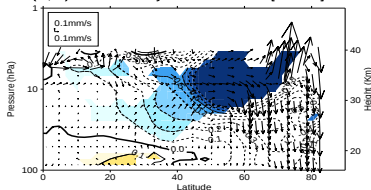
(v^* , w^*) divEP - Solar cycle Lean - Jan



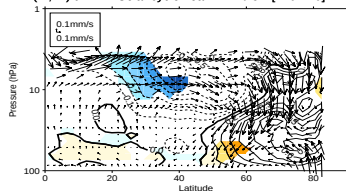
(v^* , w^*) divEP - Solar cycle Lean + 1 % UV [300-400] - Dec



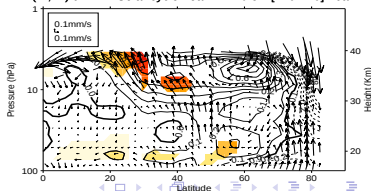
(v^* , w^*) divEP - Solar cycle Lean + 1 % UV [300-400] - Jan

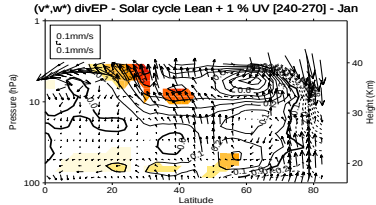
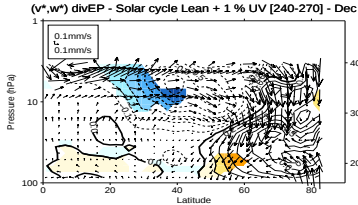
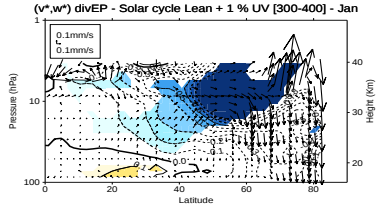
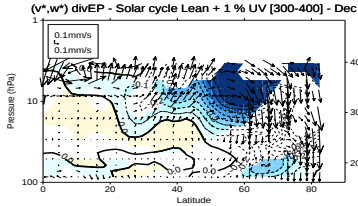
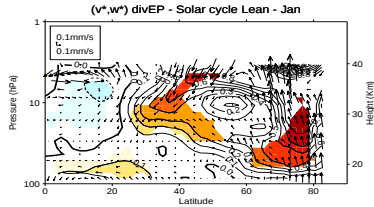
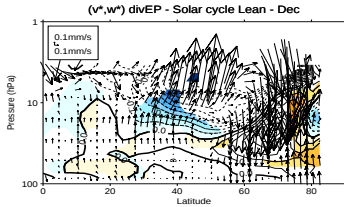


(v^* , w^*) divEP - Solar cycle Lean + 1 % UV [240-270] - Dec



(v^* , w^*) divEP - Solar cycle Lean + 1 % UV [240-270] - Jan





Solar cycle response in zonal mean zonal wind in the SH

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sensitivity to UV
forcing

Chiodo et al.

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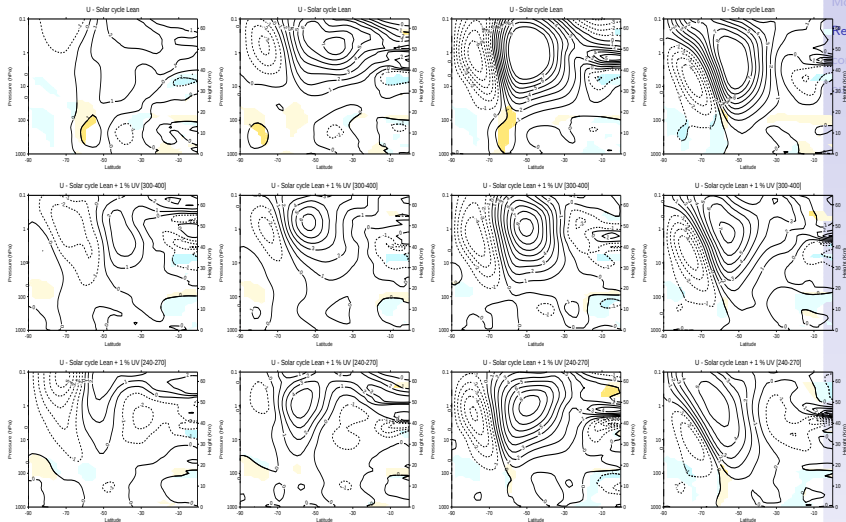
outlook

June

July

August

September

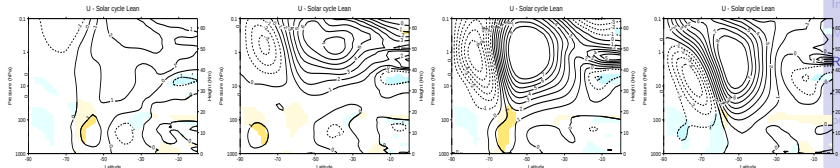


Solar cycle response in zonal mean zonal wind in the SH

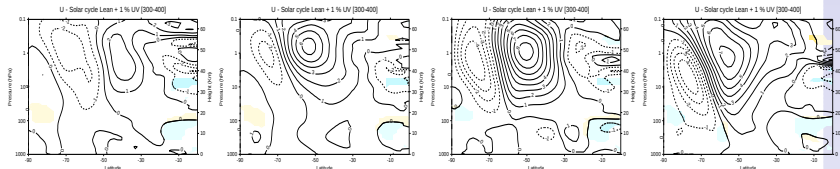
Model
sensitivity to UV
forcing

Chiodo et al.

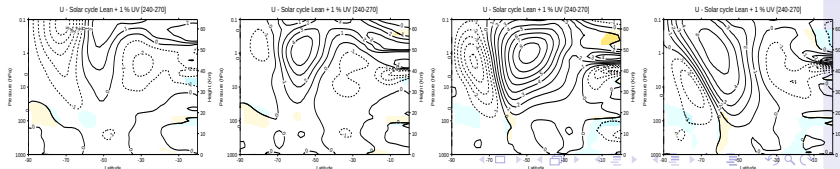
SC Lean



SC Lean + UV [300-400]



SC Lean + UV [240-270]



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- ▶ In the SH, WACCM simulates a stronger vortex during austral winter with all solar MAX spectra

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- ▶ The austral polar vortex response is stronger if the model is forced with the solar cycle variation from the Lean data-set (NRL-sMAX - NRL-sMIN)
- ▶ The reason for the difference in the SH dynamical response is still unclear

Conclusions

- ▶ The applied 1%-change in the solar radiation produces slightly different SW-heating profiles depending on the modified spectral band

Conclusions

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Conclusions

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 - ▶ Additional UV in 240-270nm leads to stronger NH polar vortex

Conclusions

- ▶ The applied 1%-change in the solar radiation produces slightly different SW-heating profiles depending on the modified spectral band
- ▶ The dynamical response in the boreal winter circulation is highly sensitive to the spectral distribution of solar cycle variations.
 - ▶ Additional UV in 240-270nm leads to stronger NH polar vortex
 - ▶ Additional UV in 300-400nm leads to more realistic signal during late-winter

Conclusions

- ▶ The applied 1%-change in the solar radiation produces slightly different SW-heating profiles depending on the modified spectral band
- ▶ The dynamical response in the boreal winter circulation is highly sensitive to the spectral distribution of solar cycle variations.
 - ▶ Additional UV in 240-270nm leads to stronger NH polar vortex
 - ▶ Additional UV in 300-400nm leads to more realistic signal during late-winter
- ▶ The response in the austral winter circulation shows less sensitivity to the idealized spectral forcing

- ▶ Investigate the tropospheric response to the spectral forcing
- ▶ Investigate the dynamical cause for hemispheric differences in the sensitivity
- ▶ Investigate the impact of spectral solar forcing on variability (Sudden Stratospheric Warmings, blocking activity etc.)